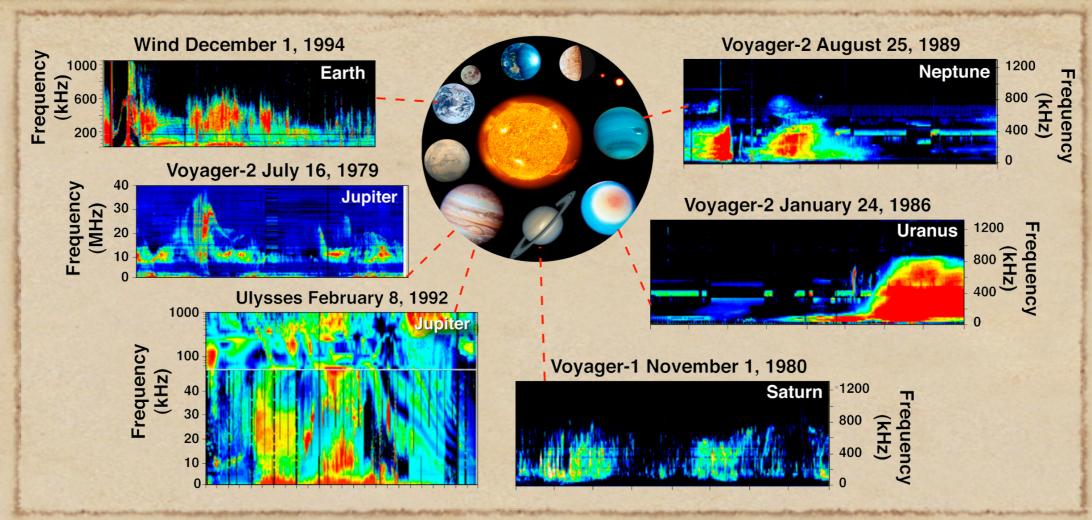
EPSC2010-394

# The Virtual Wave Observatory:

A Portal for Planetary Radio and Plasma Wave Data Leonard N. García & Shíng F. Fung NASA/Goddard Space Flight Center



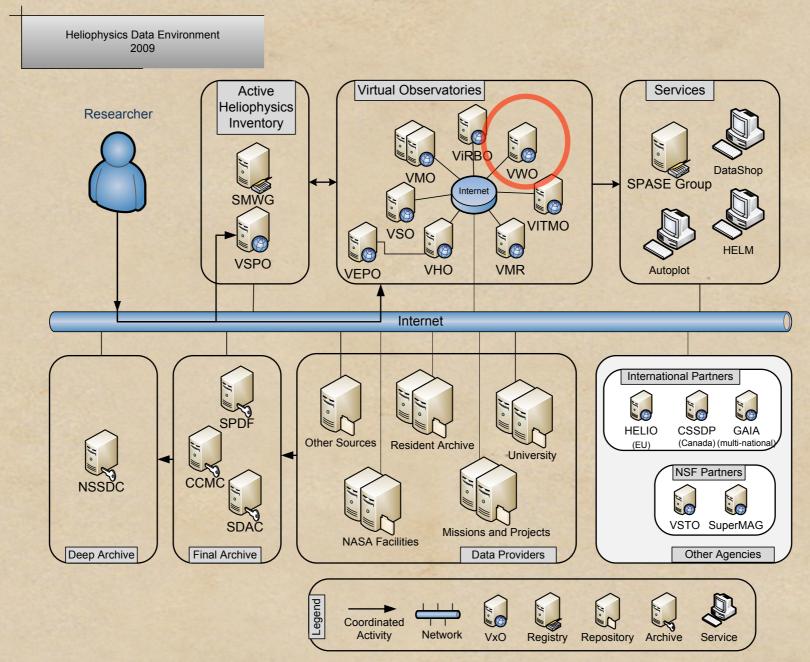
EPSC 2010, Rome, Italy 20-24 September



## Wave phenomena are ubiquitous throughout the Solar System

The goal of the Virtual Wave Observatory (*VWO*) is to make all online Heliophysics wave data searchable, usable and understandable by the scientific community.

#### VWO and NASA's Heliophysics Data Environment



Begun in 2009, VWO is one of several discipline-specific virtual observatories (VxOs) within the Heliophysics Data and Model Consortium (HDMC). NASA's HDMC oversees VxOs and Resident Archives.



Visit the Virtual Wave Observatory http://wwo.nasa.gov

#### Metadata in the SPASE Data Model

- SPASE (Space Physics Archive Search and Extract) Data Model provides common keywords and dictionaries to enable searches in distributed data archives.
- VWO works with the SPASE group to define terms for describing wave data.
- VWO is actively creating descriptions of wave data products.

```
Spase xmlns:xsi="http://www.w3.org/2001/XMLSchema-instance" xmlns="http://www.
<Version>2.0.0</Version>
<DisplayData>
     <ResourceID>spase://VWO/DisplayData/Geotail/PWI/Geotail_PWI_DS_24H</Res</pre>
    <ResourceHeader>
             <ResourceName>Geotail PWI 24 hour dynamic spectrograms /ResourceName>Geotail PWI 24 hour dynamic spectrograms 
             <AlternateName/>
             <ReleaseDate>2010-07-07T01:00:00Z</ReleaseDate>
             <Description>Geotail PWI SFA and MCA dynamic spectrogram plots
             The electric field spectrograms span the frequency range 5.62 to
             The magnetic field spectrograms also combine the MCA and SFA in
             Each plot spans 24 hours. Beneath the time axis of the magnetic
             <Acknowledgement>Before using these data in publications or pre-
             <Contact>
                     <PersonID>spase://SMWG/Person/Hiroshi.Matsumoto</person</pre>
                     <Role>PrincipalInvestigator</Role>
             </Contact>
             <Contact>
                     <PersonID>spase://SMWG/Person/Hirotsugu.Kojima</PersonIl</pre>
                     <Role>CoInvestigator</Role>
                     <Role>GeneralContact</Role>
             </Contact>
             <Contact>
                    <PersonID>spase://SMWG/Person/Guan.Le</PersonID>
                    <Role>ProjectScientist</Role>
             </Contact>
             <InformationURL>
                     <Name>Geotail PWI Instrument Page</Name>
                     <URL> http://www.kurasc.kyoto-u.ac.jp/gtlpwi/</URL>
                     <Description> Geotail PWI Instrument page maintained by
                     <Language>en</Language>
             </InformationURL>
             <Association/>
             <PriorID/>
    </ResourceHeader>
     <AccessInformation>
             <RepositoryID>spase://SMWG/Repository/RISH</RepositoryID>
             <Availability>Online</Availability>
             <AccessRights>Open</AccessRights>
             <AccessURL>
                     <Name>Geotail PWI 24 Hour Plots Data Search
                     <URL>http://www.kurasc.kyoto-u.ac.jp/gtlpwi/gtldata.htm
                     <Description/>
                     <Language>en</Language>
             </AccessURL>
             <Format>GIF</Format>
             <Encoding>None</Encoding>
             <Acknowledgement>Before using these data in publications or pre!
    <InstrumentID> spase://SMWG/Instrument/Geotail/PWI</InstrumentID>
```

#### VWO Data Registration Effort

The first Heliophysics data sets VWO is describing are those for Earth's radio and plasma wave environment.

We will also support the <u>Planetary Radio and</u> <u>Plasma Wave</u> community with data sets from:

- Voyager PRA & PWS
- Ulysses URAP
- Galileo PWS
- · Cassini RPWS

And the second of the second o	Color Sant
Satellite Mission (Experiment/Data types)	Time Span
Polar (PWI/ waveform, spectrogram)	1996- 1997
IMAGE (RPI/ spectrogram, plasmagram)	2000- 2005
Cluster (DWP, EFW, STAFF, Whisper/spectrogram)	2001- present
Cluster (WBD/ waveform)	2001- present
Geotail (PWI)	1992- present
THEMIS	2007- present
Alouette 2 (Sounder/ digital ionograms)	1965- 1975
ISIS 1 (Sounder/ digital ionograms)	1969- 1984
ISIS 2 (Sounder/ digital ionograms)	1971- 1984
DE 1 (PWI/ spectrogram)	1981- 1990
Ulysses (URAP/spectrogram, waveform, direction)	1990- 2008
Wind (Waves/ spectrogram)	1994- present
STEREO (Swaves/ spectrogram)	2006- present
Voyager 1 & 2 (PRA/spectrogram)	1977- 1989
Voyager 1 & 2 (PWS/spectrogram)	1978- 2000
Galileo (PWS/ spectrogram)	1989- 2003
Cassini (RPWS/spectrogram)	1997- present
Hawkeye (ELF- VLF/spectrogram)	1974- 1978
ISEE 1 & 2 (Plasma wave, VLF)	1977- 1987
ISEE 3/ICE (Plasma wave spectrum analyzer)	1978- 1997
ISEE 3/ICE (Radio mapping of Solar Wind disturbances)	1978- 1997
CRRES (Plasma wave)	1990- 1991

Ground-based (Experiment/Data types)	Time Span
Augsburg College AGO (Searchcoil Magnetometer /spectrograms)	1993- 2008
U. Maryland AGO (Magnetometer, Riometer & VLF /Survey plots)	1986- 2005
Augsburg College Svalbard (Searchcoil Magnetometer /spectrograms)	2006- 2008
NCAR Svalbard (Searchcoil Magnetometer /daily data files)	2006- 2008

#### Searches by Context

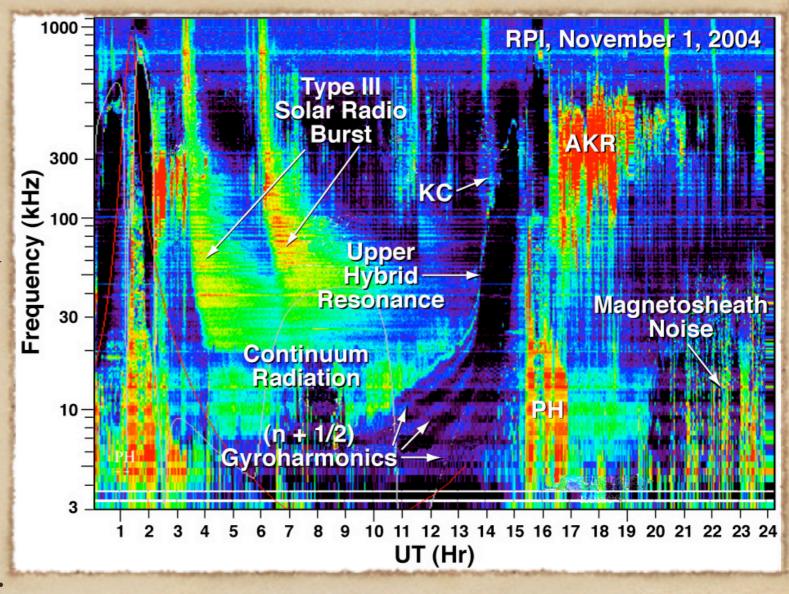
- Natural wave emissions in the Earth's magnetosphere are often dependent on the state (i.e. context) of the magnetosphere.
- VWO has enabled searches on time-shifted solar wind, IMF and geomagnetic conditions.
- As an example, correlations have been shown between Saturn SKR and solar wind (Desch and Rucker, 1983).

#### Searches by Phenomenon

• In order to help (Under Development)

broaden the use of wave data to non-wave experts, VWO intends to develop a service to capture data annotations provided by wave experts as they analyze the data

 These annotations will then be stored in a searchable database to enable searches by these wave phenomena.



Success of this service depends on community input.

#### Cooperation with other VxOs

- The VWO uses the SPASE-Query Language (SPASE-QL, http://spaseql.gsfc.nasa.gov) to support seamless and transparent interactions with other VxOs to obtain matches to user search queries.
- This communication method will also allow users to access ancillary data not served by the VWO (for example particle and DC field data or models).

### Interacting with the VWO

- The VWO is meeting with the wave research community to ensure that our search tools respond to their needs.
- The VWO is working with data providers to ensure accuracy and completeness of data descriptions.
- The VWO web site will accept feedback from users.
- The VWO will depend on community involvement for development of the annotation service.

#### Additional Information

• Fung, S. F., The Virtual Wave Observatory (VWO): A Portal to Heliophysics Wave Data, Radio Science Bulletin, No. 332, pp. 89-102, March 2010.

#### **VWO Contacts:**

- Shing.F.Fung@nasa.gov
- Leonard.N.Garcia@nasa.gov